



MICROSCOPE data processing and analysis, missing data and inpainting

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with

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return on innovation

Data analysis: what do we measure?

EP violation signal

$$\delta = \frac{m_{g2}}{m_{I2}} - \frac{m_{g1}}{m_{I1}}$$

Earth gravity gradient tensor: *model, S/C position from SST and instr. data and attitude*

Instrument's moments of inertia: *minimized by AOCS model, S/C position from SST and instr. data and attitude*

Separation between test masses' centers: *minimized by instrument design*

Coriolis acceleration

Relative acceleration

Non-gravitational accelerations => *drag-free system*

$$\vec{\Gamma}_{\text{meas,d}} = [\mathcal{M}_c] \left(\delta \vec{g} + ([\mathcal{T}] - [\mathcal{I}n]) \vec{\Delta} - 2 [\Omega] \dot{\vec{\Delta}} - \ddot{\vec{\Delta}} \right) + \vec{K}_{0,d} + [\mathcal{M}_d] \left(\vec{\Gamma}_{\text{app,c}} + \frac{\vec{F}_{\text{ext}}}{M_{\text{Isat}}} + \frac{\vec{F}_{\text{th}}}{M_{\text{Isat}}} \right) + \frac{1}{2} K_{2,i} \Gamma_{\text{App},i}^2 - \frac{1}{2} K_{2,j} \Gamma_{\text{App},j}^2 + \vec{\Gamma}_{\text{n,d}} + [\mathcal{C}_d] \dot{\vec{\Omega}}$$

Bias difference: *limited thermal fluctuations*

Differential mode sensitivity matrix: *estimated by in-flight calibration*

Common mode sensitivity matrix: *estimated by in-flight calibration, limited by design*

$$\begin{pmatrix} K_{cx} & \eta_{cz} + \theta_{cz} & \eta_{cy} - \theta_{cy} \\ \eta_{cz} - \theta_{cz} & K_{cy} & \eta_{cx} + \theta_{cx} \\ \eta_{cy} + \theta_{cy} & \eta_{cx} - \theta_{cx} & K_{cz} \end{pmatrix}$$

Scale factor

Coupling between instrument's axes

Instrument's axes / satellite reference frame alignment

Quadratic factor

Total acceleration applied to proof-mass

Angular-linear couplings

Instrument noise

How to extract the information?

Least-Square estimation after correction / modelling of nuisance parameters

$$\begin{aligned} 2\Gamma_x^{(d)} &\simeq 2Bd_x \\ &+ ac_{11} [\delta g_x + (T_{xx} - In_{xx})\Delta_x + (T_{xy} - In_{xy})\Delta_y + (T_{xz} - In_{xz})\Delta_z] \\ &+ ac_{12} [\delta g_y + (T_{yx} - In_{yx})\Delta_x + (T_{yy} - In_{yy})\Delta_y + (T_{yz} - In_{yz})\Delta_z] \\ &+ ac_{13} [\delta g_z + (T_{zx} - In_{zx})\Delta_x + (T_{zy} - In_{zy})\Delta_y + (T_{zz} - In_{zz})\Delta_z] \\ &+ ac_{11} (2\Omega_z \dot{\Delta}'_y - 2\Omega_y \dot{\Delta}'_z - \ddot{\Delta}'_x) \\ &- 2\theta_{c,x} (\Omega_y \dot{\Delta}'_y + \Omega_z \dot{\Delta}'_z) - 2ac_{13} \Omega_x \dot{\Delta}'_y + 2ac_{12} \Omega_x \dot{\Delta}'_z \\ &+ \frac{K_{21,xx}}{K_{1,x}^2} (\Gamma_x^{(1)} - B1_x)^2 - \frac{K_{22,xx}}{K_{2,x}^2} (\Gamma_x^{(2)} - B2_x)^2 \\ &+ 2ad_{11} (\Gamma_x^{(c)} - Bc_x) + 2ad_{12} (\Gamma_y^{(c)} - Bc_y) + 2ad_{13} (\Gamma_z^{(c)} - Bc_z) \\ &+ 2nd_{11} \dot{\Omega}_x + 2nd_{12} \dot{\Omega}_y + 2nd_{13} \dot{\Omega}_z \end{aligned}$$

Estimated through calibration

Measured and/or computed

Negligible at f_{EP}

Slide courtesy G. Métris

Instrument's noise spectral density

Touboul 2009

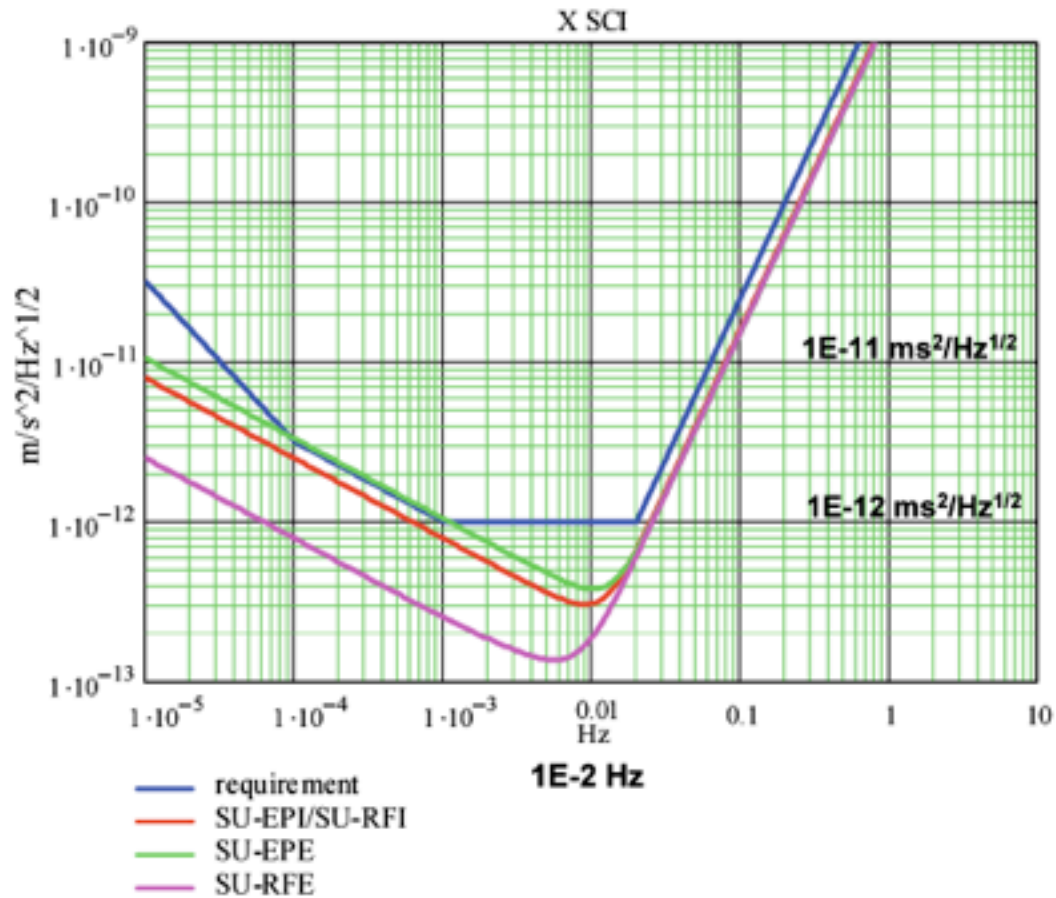
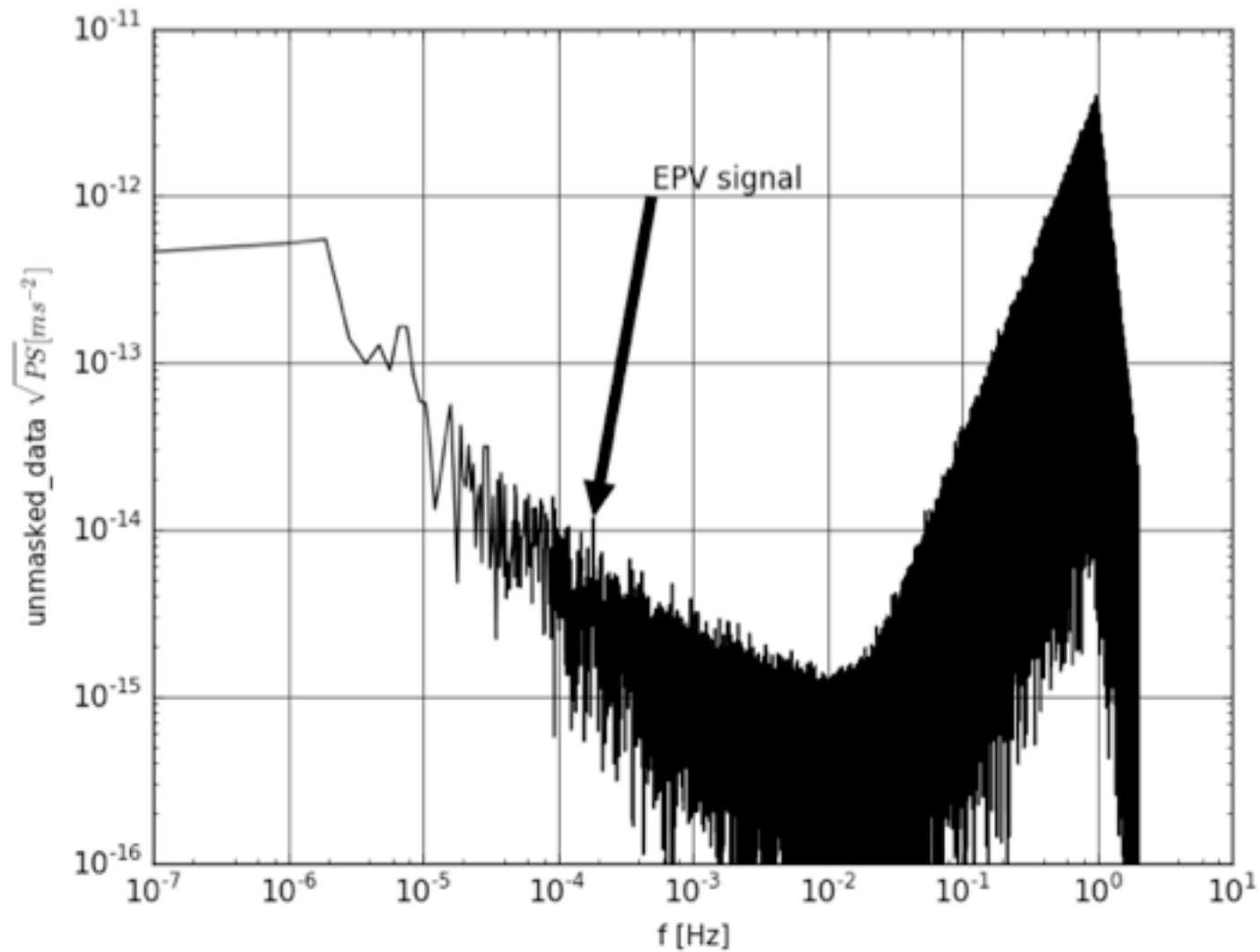


Fig. 4 $(PSD)^{1/2}$ expressed in $\text{ms}^{-2} \text{Hz}^{-1/2}$ of all inertial sensors: Platinum external mass (in purple), Platinum internal mass (in red), Titanium external mass (in green), objectives to be reached (in blue)

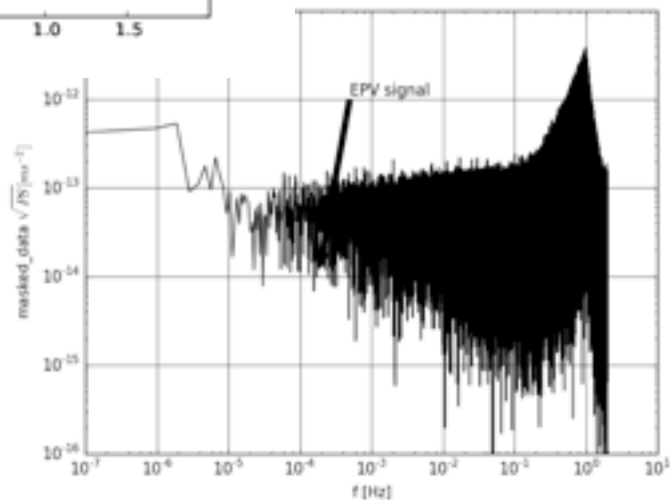
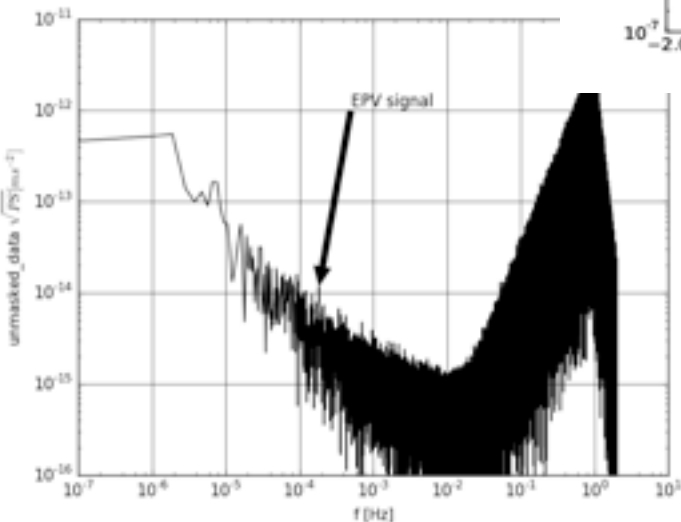
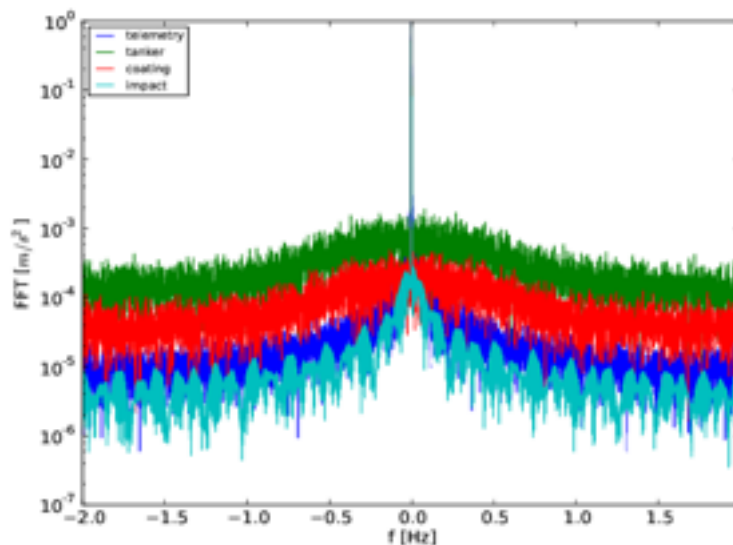
With nuisance parameters perfectly calibrated / corrected for...



- **Teletransmission errors**
 - Information from Picard mission:
 - **frequency**: about 100 events over 10 months
 - **duration**: from seconds to hours
- **Coating cracking**
 - due to temperature changes (Earth / Space vacuum)
 - **frequency**: for each of the four satellite sides, about 6 times when the side faces the Earth
 - **duration**: 0.5-0.75s → 2 - 3 measurement points
- **Tank cracking**
 - worst case, depending on gas pressure
 - **frequency**: for each of the 6 tanks, about 43 times/orbit
 - **duration**: 0.5s → 2 measurement points

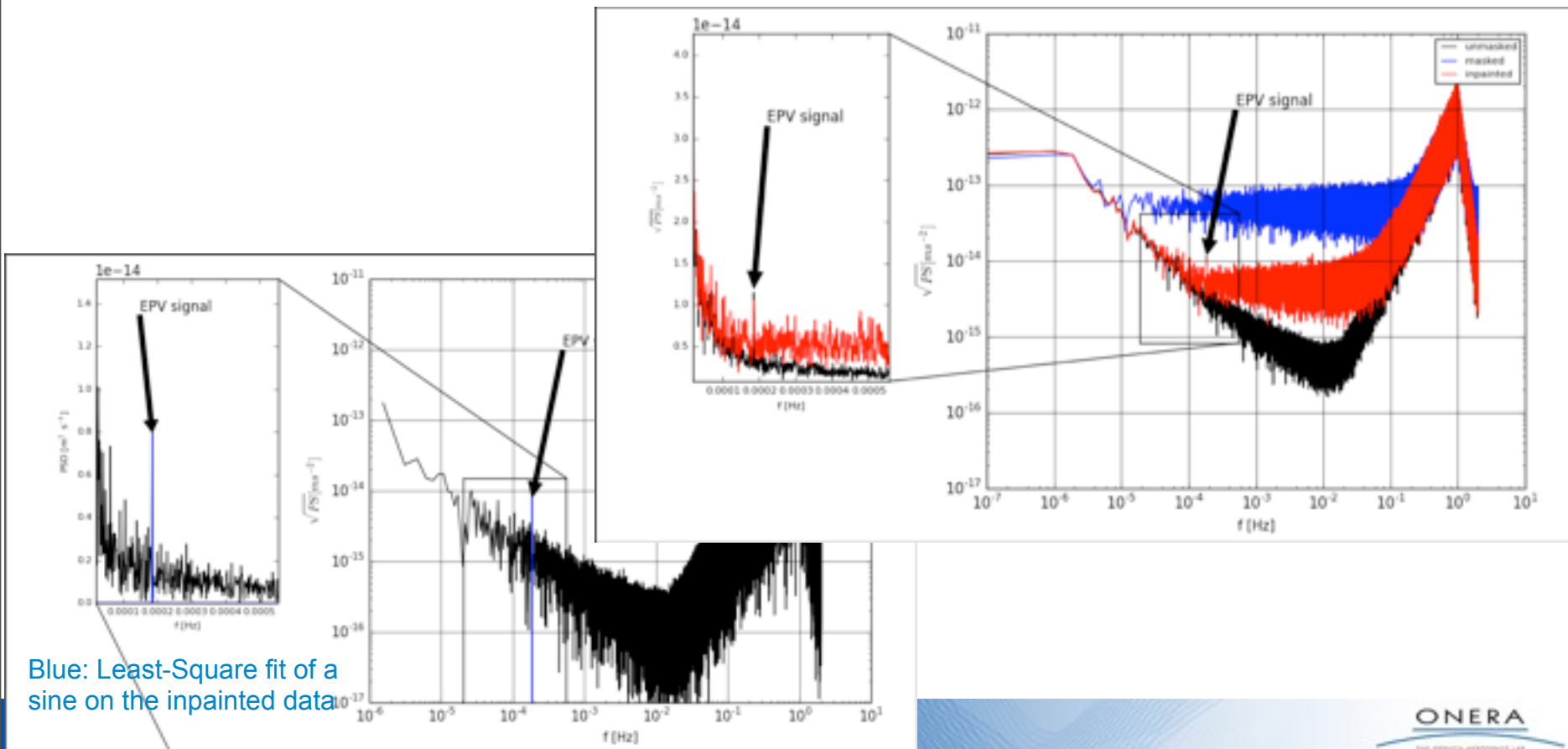
Impact of missing data

Convolution of non-trivial spectral window with (signal+noise)'s power spectrum
=> **spectral leakage**

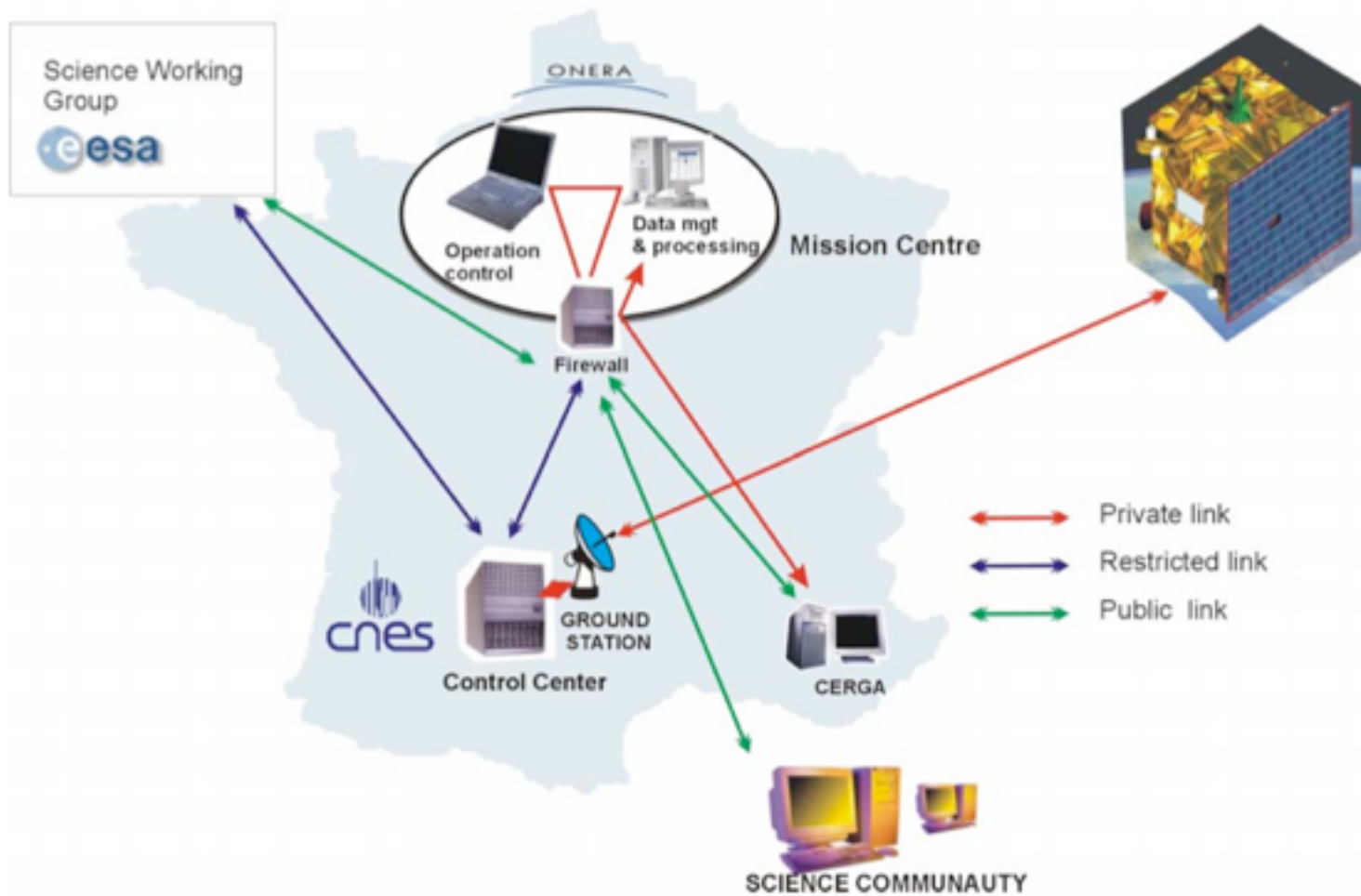


Inpainting application to MICROSCOPE

- ✓ Inpainting developed by AIM/CEA, originally for 2D astronomy images; already adapted to time series for asteroseismology (J.L. Starck, S. Pires, S. Mathur's talks)
- ✓ Extrapolation of the missing information with sparsity prior on the solution: there is a function dictionary on which the complete data are sparse (very few non-negligible coefficients) but the incomplete data are not sparse

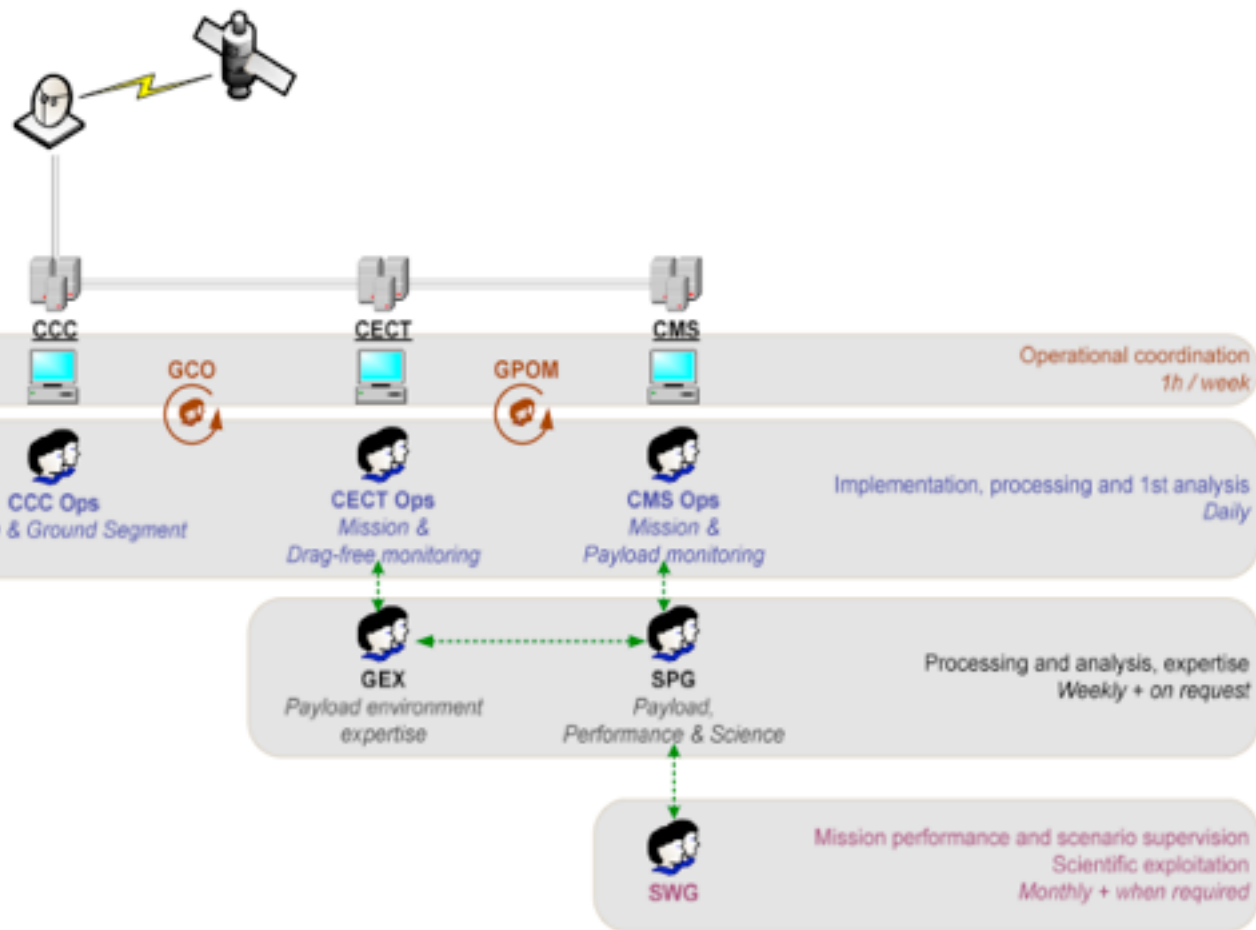


MICROSCOPE ground segment



Ground segment organization

3 levels



- permanent activity for data processing
- monthly meetings
- weekly potential request for mission scenario & operation

- biannual meetings or quarterly for data processing organization and validation
- monthly potential requests for mission scenario

MICROSCOPE Science Mission Center (CMSM)

- Ensure all operational functions to maximize instrument's operation
- Day-to-day instrument management and monitoring
- Weekly mission performance check
- Propose modifications to the mission scenario
- Data release and archiving

→ **Different time scales with different flexibilities:**

• **1-week horizon: operational loop**

- Verification of data integrity by automated processing
- Mission program fixed (except potential stop or extension of a long session)

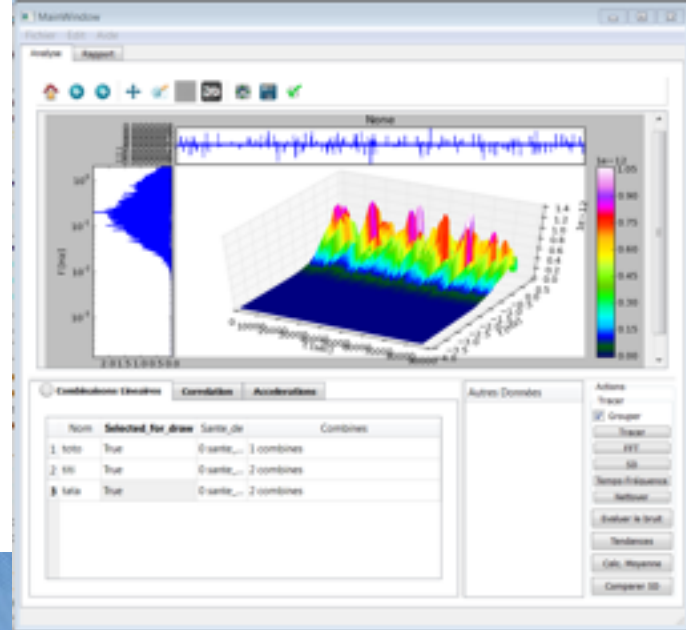
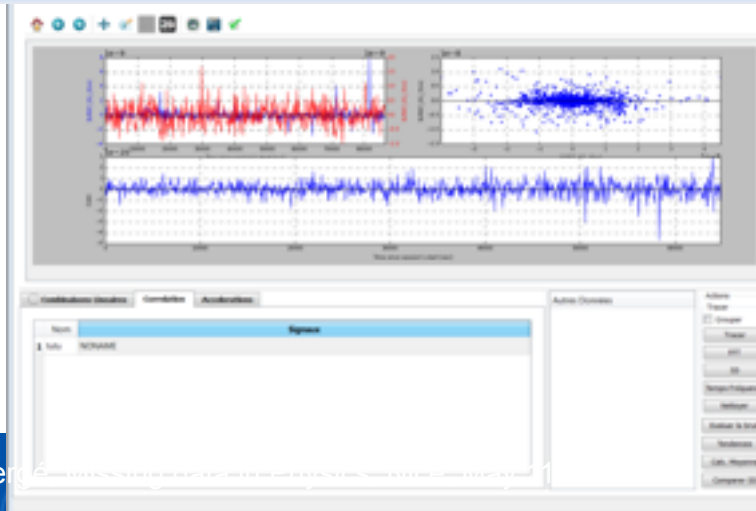
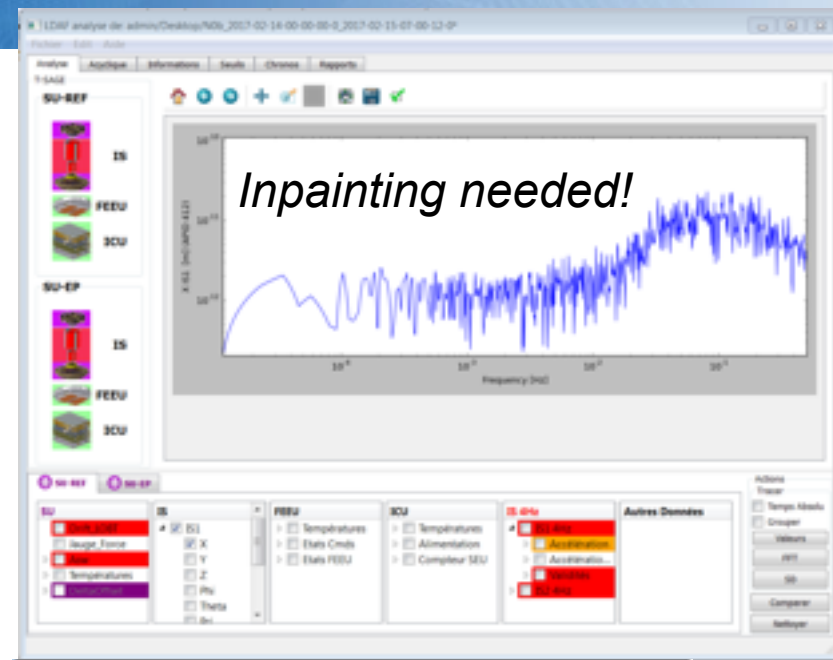
• **1-month horizon:**

- Preliminary analysis of data
- Scenario still modifiable, in the frame of the predefined sessions

• **1-year horizon:**

- Detailed scientific analysis
- Detailed performance analysis
- Optimization of calibration processing
- Application of data correction models

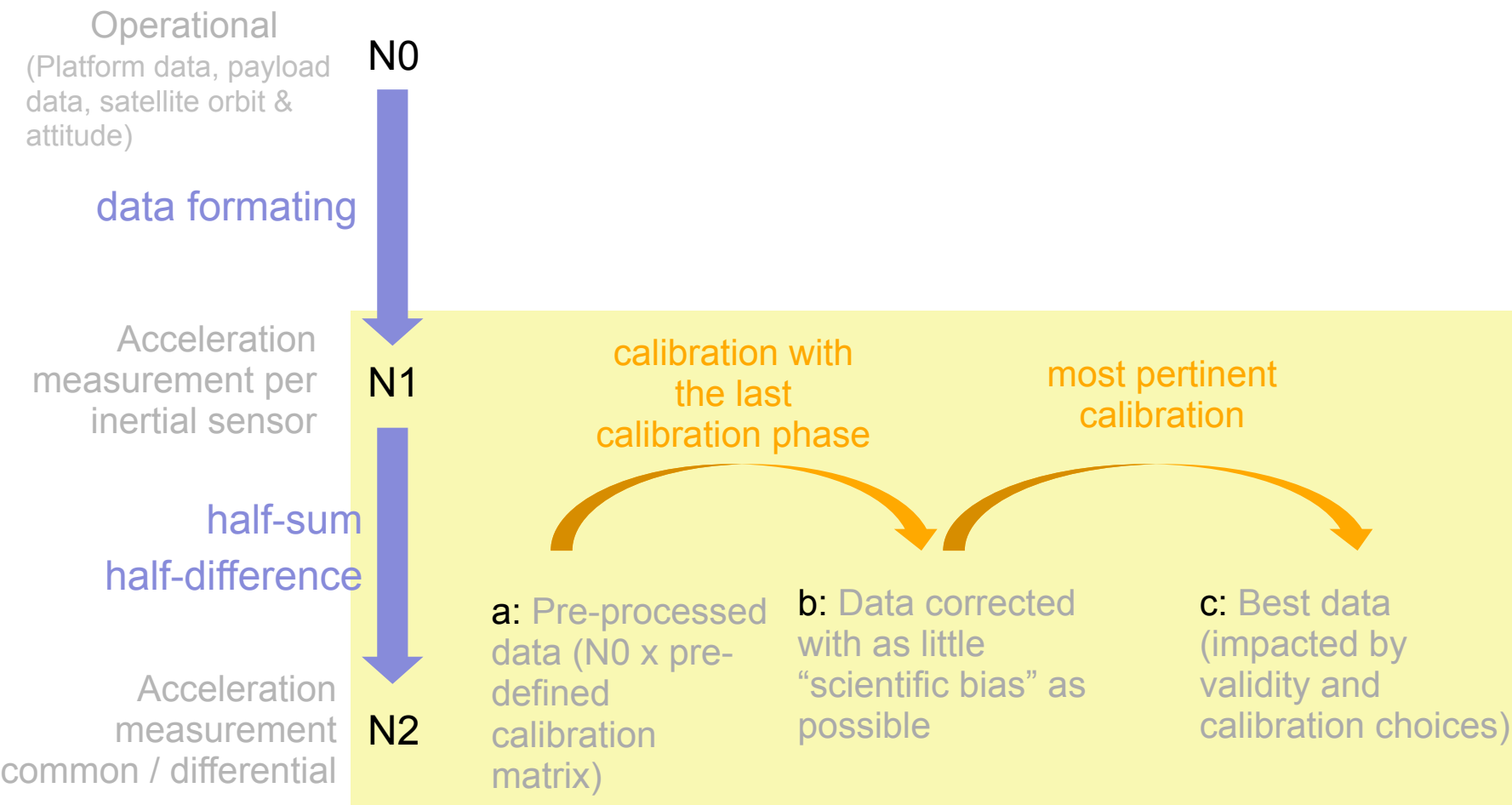
Daily / Weekly operations: instrument monitoring



Weekly / monthly: data analysis

- ★ Data architecture organization (see next slide)
- ★ Calibration: to go from raw data to usable data, usually on differential accelerations
 - from on ground controls and models
 - from in-flight assessments
 - from in-flight dedicated sessions
 - Most obvious calibrations: gravity gradient tensor
 - Different levels of calibration
- ★ Treatment of invalid / missing data
 - detection according to instrumental flags (provided in TMs) and/or thresholds crossings -> creation of a mask
 - (optional) inpainting
- ★ Computation of differential accelerations (which are used to look for Equivalence Principle violation)
- ★ Search and estimation of the Equivalence Principle violation (Least-Square, KARMA -- see Q. Baghi's talks)
- ★ Data archiving and distribution

MICROSCOPE data architecture: 3 levels



Conclusion and prospects

- ★ MICROSCOPE data analysis is a complex task: look for and estimate a very weak signal in a colored-noise dominated time-series

- ★ Made much more difficult by missing data: leakage of high-frequency noise on frequencies where the signal is expected
- ★ Inpainting allows us to correct for missing data well enough to detect the signal
- ★ Inpainting will be used from day-to-day instrument's monitoring operations to scientific expertise and analysis

- ★ CMSM is responsible for the data processing, analysis, archiving and distribution
- ★ Different levels of data will be provided:
 - ◆ Level 1 data (organized by inertial sensor) can be used to:
 - ✓ estimate Eötvös parameter from data weakly impacted by CMSM's validity and calibration choices
 - ✓ perform other applications (gravity, aeronomy...) not depending on differential mode
 - ◆ Level 2 data (optimized for the differential mode)
 - ✓ include corrections
 - ✓ allow most direct estimation of the Eötvös parameter
 - ◆ Auxiliary data: orbit, attitude...